Madhyamānayanaprakāraḥ: a hitherto unknown manuscript ascribed to Mādhava

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Abstract

It is generally believed that only two works of Mādhava, namely Venvāroha and Sphutacandrāpti are extant today. However, it seems a few of his small tracts have also survived. The present article deals with one such small tract entitled 'Madhyamānayanaprakārah' ('the procedure for obtaining the mean') found among the K. V. Sarma collection of manuscripts. Starting with a graphic description of the geometrical construction involved in the manda-samskāra, the manuscript proceeds to describe in great detail the concept of *viparītakarņa*. Along with the formula for the *viparītakarna*, it also presents a detailed discussion on its application to bring about a simplification—that enables an astronomer to avoid the strenuous and persevering task of performing an iterative process—in the computation of aviśista-manda-karna. Finally it concludes with the description of the procedure for the computation of the madhyamagraha (mean longitude of a planet) from the manda-sphuta-graha (longitude of the planet corrected for eccentricity), which indeed explains the choice of the title Madhyamānayanaprakāra.

Keywords: madhyama, manda-sphuța, viparīta-karņa, aviśiṣța-manda-karņa

1 Introduction

The Kerala school of astronomy pioneered by Mādhava (c. 1340-1420) of Sangamagrāma seems to have blazed a trial in the study of a branch of mathematics

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that goes by the name of analysis today. His major contributions include not only the discovery of the infinite series for π , but also several fast convergent approximations to it. Unfortunately, we do not have any manuscript attributed to Mādhava that contains verses presenting the above series. It is only through the quotations and citations made the successors of Mādhava, we come to know that it was Mādhava who discovered the infinite series for π as well as the power series for the sine and cosine functions, which were re-invented in the West three centuries later.

Mādhava's contributions to astronomy are not as well known as his contributions to mathematics. It is generally believed that only two works of him, namely Venvaroha and $Sphutacandrapti^1$ —both related to the computation of Moon's positions—are extant today. However, K. V. Sarma ascribes one more work on astronomy entitled Aganitagrahacara to Mādhava, which is yet to see the light of the day. Besides editing and publishing several manuscripts, the renowned indologist Sarma, has also painstakingly collected paper transcripts of a plenty of palmleaf manuscripts, that are in possession of various oriental manuscript libraries.

All these transcripts, as well as his personal collection of thousands of books are currently preserved in Sree Sarada Education Society and Research Centre.² Among this collection, recently we came across a couple of short tracts in Sanskrit prose that have been identified by Sarma as due to Mādhava. One such tract entitled '*Madhyamānayanaprakāraḥ*' which literally means 'the procedure for obtaining the mean', has been chosen for study in the present article. Here, besides explaining the technical content of the manuscript in the light of modern mathematical language, we also present an edited version of the text along with English translation.

The article is divided into six sections including the present one on introduction. Sections 2 and 3 present the details of the manuscript material employed and the scheme adopted in preparing an edited version of the text. While section 4 is devoted to explain the content of the $m\bar{u}la$ -slokas (the original verses on which a commentary is being authored) using modern mathematical notations, section 5 presents the edited text along with English translation. In the last and the final section we give detailed explanatory notes on the manuscript material

¹These two texts have been critically edited and published by K. V. Sarma.

²An institution that was founded by Sarma himself, by way of donating his own savings as well as invaluable collection of books and manuscripts, to foster indological studies. The address of this institution, currently renamed as K. V. Sarma Research Foundation is: 2, East Coast Flat, 32, II Main Road, III Cross, Gandhi Nagar, Adyar, Chennai 600 020.

by dividing it into various subsections that seemed convenient and appropriate.

2 The Manuscript material

The edition of the text *Madhyamānayanaprakāra* presented below is based on the lonely paper transcript of the orginial palmleaf manuscript in the possession of Indian Office Library, London. The paper transcript, in four pages (19–22) found among the K. V. Sarma manuscript collection and bears the number **KVS 354-C**. It may be noted on the first page of the transcript (see Figure 1), that in the top-left-corner there is a marking that mentions the number **IO 6301, Ms. # 16–18**.

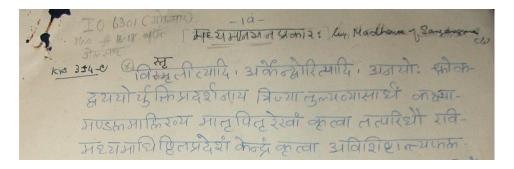


Figure 1: The first few lines of the first page of the manuscript

From this noting (in pencil) made by Sarma, we understand that the original manuscript is with the Library of India Office, London, and that this material is in three folios (16–18). More details about this manuscript have been provided by A. B. Keith in the catalogue prepared by him³ wherein it has been mentioned that the manuscript is on $8\frac{3}{8}'' \times 1\frac{5}{8}''$ talipat (palmleaves) and that it in Malayalam script.

3 Editorial note

While studying the manuscript (**KVS 354-C**), which is in Devanāgarī script as seen from Figure 1, it was quite evident to us that the readings in certain

³A. B. Keith, Catalogue of the Sanskrit and Prakrit manuscripts in the Library of the India Office, vol. 2, Oxford 1935, pp. 774–5, (no. 6301).

places were definitely fautly. As an illustration we present a couple of examples. In one instance, a string of words was simply repeating (see Figure 2).

मु गक्म दानूण धर्त च कत्वा वरी त्रिज्या लभ्यते। नगे स्फुटनोटिफा मुंगनन्या कावणे हार्तन क्राता तरी त्रिज्या कर्म्यते , कोई स्प्रुटकोटि फर्क मुरा-नक्यीयानू गे राजे - न क्राता नरी त्रिल्या लभ्यते.

Figure 2: Clip of the manuscript where a string of words get repeated.

As it is pretty obvious that this is due to scribal error, in our version of the edited text this was dropped out. In yet another place (see the fourth line of the clip of the manuscript shown in Figure 3, whose typed version is given below), it was noted that a string of characters were missing from a compound word.

तदर्धं मन्दोचोनमा ल्लब्धं अविशिष्टभुजाफलम्।

With reasonable working knowledge in Sanskrit it should not be difficult to make out that there is obviously some error in the second and the third words though fixing the error requires the knowledge of the subject as well as intelligent guess work. After pondering over the sentence, keeping the context in mind, it became evident to us that the reading should be

तदर्धं मन्दोचोनमध्यमाल्लव्यं अविशिष्टभुजाफलम्।

रद्रोण द्वितीयभष्यु खनी सवसमाहिरवेत् पूतः -पूर्विति खिली चारीन परिसे नजीवन परिदोद्य सेपातक्व प्रापिणे स्रामिश्वितः प्राः पूर्वकिलिनिद्यतीय तिद्धीं मन्दी चोज मा क्रूथ्मं आविश्विष्ठमुजापलम्

Figure 3: Clip of the manuscript where a few letters are found to be missing.

While presenting an edited version of the text, to the best or our ability we have tried to fix the errors and present a correct version of the original. In doing so, the reading found in the manuscript had to emended at places. Though we have been very careful and judicious in emending the text, for the benefit of scholars to judge for themselves, we have presented the readings found in the manuscript in the form of footnotes with the exception of one type of emendation. On a few places the $dandas^4$ in the manuscript were found to be misplaced and hence had to be removed and inserted in certain other places. Such insertion and deletion of dandas have been done silently, where it seemed appropriate to us.

Before we present the edited text and the translation, we offer a brief explanation of the verses that form the $m\bar{u}la^5$ of the manuscript and also discuss on the authorship of the same.

4 The $m\bar{u}la$ verses

The manuscript essentially presents a detailed commentary on two verses commencing with '*vistrti*' and '*arkendu*'. In fact, this has been clearly stated in the very opening sentence of the manuscript:

'vistṛtī'tyādi| 'arkendvor'ityādi| anayoḥ ślokadvayayoḥ yuktipradarśanāya ...

[The verse] that commences with *vistrti*. [The verse] that commences with *arkendvoh*. In order to present the rationale behind the [content] of these two verses, ...

Though these two verses happen to be part of the second chapter of the celebrated work *Tantrasangraha* of Nīlakanṭha Somayājī, it turns out that Nīlakanṭha is not the author of them. Neither is there a mention about the author of the $m\bar{u}la$ verses anywhere in the text of the manuscript itself. However, the quotations made by Nīlakanṭha—with due acknowledgement to the authors whom he is quoting—in his works *Tantrasangraha*,⁶ Siddhāntadarpaṇa and Āryabhaṭīyabhāsya enable us to identify the author of the $m\bar{u}la$ verses to be Mādhava.

⁴This term refers to the vertical bars employed to indicate the end of a sentence particularly while using a $Devan\bar{a}gar\bar{i}$ script.

⁵The world $m\bar{u}la$ which literally the 'root' or 'basis', is traditionally employed to refer to the text on which commentary is being written.

⁶See for instance, $\{TS \ 1958\}, (II.44).$

4.1 About the authorship

Nīlakantha in his *Tantrasangraha*,⁷ soon after quoting the verse *vistrti* states:

इति वा कर्णः साध्यः मान्दे सकृदेव माधवप्रोक्तः।

Or, the *karṇa* is to be obtained only once [in this way] in the *manda* process as enunciated by Mādhava.

In the $Ny\bar{a}yabh\bar{a}ga$ of his $Siddh\bar{a}nta-darpana-vy\bar{a}khy\bar{a}^8$ (auto-commentary on $Siddh\bar{a}nta-darpana$), in the context of explaining how the problem of mutual dependency involved in the computation of karna and the radius of the manda-virta can be circumvented, Nīlakantha observes:

स्ववृत्तकलाप्रमितस्य कर्णस्य सदैव त्रिज्यातुल्यत्वात् तेन मध्यकक्ष्याव्यासार्धा-नयने कर्णानयनविपरीतकर्म कार्यम्। तद्य माधवेनोक्तम् – विस्तृतिदल...

Since the karna measured in terms of the minutes of its own circle is [taken to be] equal to $trijy\bar{a}$, in order to compute the radius of the kakṣyāmaṇdala (the deferent circle), we have to adopt a process that is [exactly] the reverse of the process employed in finding the karṇa. And that has been stated by Mādhava [thus] – vistṛti ...

In his magnum opus $\bar{A}ryabhatiya-bhaiya$, while presenting a detailed commentary (that runs to more than 20 pages) on the five verses⁹ of $\bar{A}ryabhati that describe the geometrical picture of planetary motion, Nīlakantha says:$

स्फुटे(न) मध्यमानयने सकृत्कर्म अन्यादृशं माधवोक्तमपि श्रुतम् – 'अर्केन्द्रोः…'

In obtaining the mean from the true, the one-step process enunciated by Mādhava is also heard - '*arkendvo*<u>h</u>...'

These acknowledgements made by Nīlakantha clearly settle the issue regarding the authorship of the $m\bar{u}la$ verses and leave no room for ambiguity, whatsoever. This, however, is not the case with respect to the prose commentary in the manuscript. Our ascription of it to Mādhava is solely based on the noting made by Sarma—just next to the title (see Figure 1).

⁷See **{TS 2010}**, p. 105.

⁸{SDA 1976}, p. ...

⁹Āryabhatīya, Kālakriyāpāda, verses 17–21.

4.2 Explanation of the verse 'vistrti ...'

The verse commencing with '*visiți*' presents the formula enunciated by Mādhava for finding the *aviśiṣṭa-manda-karṇa*—the iterated hypotenuse associated with the *manda-saṃskāra*. At this stage, it would be useful to include a brief note on *aviśiṣṭa-manda-karṇa* as well as procedure for the computation of the *manda-sphuṭa*¹⁰ in order to make the discussion more edifying. This would also enable the reader to have a better appreciation of the substantial simplification in computation achieved by Mādhava in obtaining *aviśiṣṭa-manda-karṇa* without having to resort to the conventional interative process.

4.2.1 Calculation of manda-sphuta

The procedure for obtaining the manda-sphuța from the mean longitude (madhyama) of the planet prescribed in the Indian astronomical texts can be understood with the help of Figure 4. Here, $\theta_0 = \Gamma \hat{O} P_0$ represents the longitude of the mean planet (madhyama-graha, P_0), $\theta_0 = \Gamma \hat{O} U$ the longitude of mandocca (apogee or aphelion) and $\theta_{ms} = \Gamma \hat{O} P$ the longitude of the manda-sphuța which is to be determined from θ_0 . It can be easily seen that

$$\begin{aligned}
\theta_{ms} &= \Gamma OP \\
&= \Gamma \hat{O} P_0 - P \hat{O} P_0 \\
&= \theta_0 - \Delta \theta.
\end{aligned}$$
(1)

Since the mean longitude of the planet θ_0 is known, the manda-sphuta θ_{ms} is obtained by simply subtracting $\Delta \theta$ from the madhyama. In the figure, $P_0P = r$ and $OP_0 = R$ represent the radii of the epicycle and deferent circle respectively. Considering the right-angled triangle OPQ, we have

$$K = OP = \sqrt{OQ^2 + OP^2} = \sqrt{(OP_0 + P_0Q)^2 + OP^2} = \sqrt{\{R + r\cos(\theta_0 - \theta_m)\}^2 + r^2\sin^2(\theta_0 - \theta_m)}.$$
 (2)

Also

$$K\sin\Delta\theta = PQ$$

= $r\sin(\theta_0 - \theta_m).$ (3)

¹⁰The longitude of the planet obtained by applying the *manda-saṃskāra* (equation of centre) to the mean longitude of the planet.

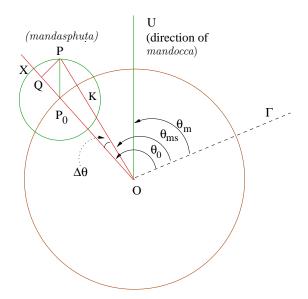


Figure 4: Geometrical construction underlying the rule for obtaining the *manda-sphuţa* from the *madhyama* using the epicycle approach.

Multiplying the above by R and dividing by K we have

$$R\sin\Delta\theta = \frac{r}{K}R\sin(\theta_0 - \theta_m). \tag{4}$$

According to the geometrical picture of planetary motion given by Bhāskara I (c. 629), the radius of the epicycle manda-nīcocca-vrtta (r) employed in the the manda process is not a constant. It varies continuously in consonance with the hypotenuse, the manda-karṇa (K), in such a way that their ratio is always maintained constant and is equal to the ratio of the mean epicycle radius (r_0) —whose value is specified in the texts—to the radius of the deferent circle (R). Thus, according to Bhāskara, as far as the manda process is concerned, the motion of the planet on the epicycle is such that the following equation is always satisfied:

$$\frac{r}{K} = \frac{r_0}{R}.$$
(5)

Thus the relation (4) reduces to

$$R\sin\Delta\theta = \frac{r_0}{R} R\sin(\theta_0 - \theta_m), \qquad (6)$$

where r_0 is the mean or tabulated value of the radius of the manda epicycle.

4.2.2 Bhāskara's method for obtaining the avisista-manda-karna

It may be noted that in (5), RHS is the ratio of two fixed values, namely the mean epicycle radius and the $trijy\bar{a}$. This, however is not the case with respect to quantities appearing in the LHS. Here, both the numerator and the denominator are variables. Hence the question arises, as to how one can obtain the values of r and K at any given instant, though the ratio is always a constant. For this, Bhāskara prescribes an iterative procedure called *asakrt-karma*, by which both are simultaneously obtained.

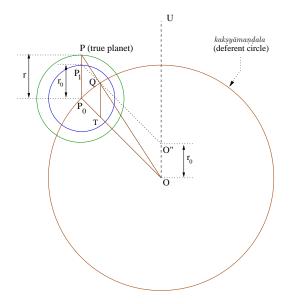


Figure 5: The variation of the radius of the manda epicycle with the mandakarna.

We explain this with the help of Figure 5. Here P_0 represents the mean planet around which an epicycle of radius r_0 is drawn. The point P_1 on the epicycle is chosen such that PP_1 is parallel to the direction of the mandocca, OU. Now, the first hypotenuse (sakrt-karna) is found from r_0 using the relation

$$OP_1 = K_1 = [(R\sin(\theta_0 - \theta_m))^2 + (R\cos(\theta_0 - \theta_m) + r_0)^2]^{\frac{1}{2}}.$$
 (7)

From K_1 , using (5), we get the next approximation to the radius $r_1 = \frac{r_0}{R}K_1$. Then, from r_1 we get the next approximation to the karna,

$$K_2 = [\{R\sin(\theta_0 - \theta_m)\}^2 + \{R\cos(\theta_0 - \theta_m) + r_1\}^2]^{\frac{1}{2}},$$
(8)

and from that we get $r_2 = \frac{r_0}{R}K_2$. The process is repeated till the radii and the *karnas* do not change (*aviśesa*). The term *aviśesa* means 'not distinct'. In

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the present context it means that the successive karnas are not distinct from each other. That is, $K_{i+1} \approx K_i = K$. If this is satisfied, then $r_{i+1} \approx r_i = r$. Consequently, the equation giving the manda-correction (4) becomes

$$R\sin\Delta\theta = \frac{r}{K}R\sin(\theta_0 - \theta_m) = \frac{r_0}{R}R\sin(\theta_0 - \theta_m).$$
(9)

Thus the computation of $\Delta\theta$ —known as the mandaphala, that is to be applied to the madhyama to get the manda-sphuta—involves only the mean epicycle radius and the value of the trijyā and not the value of the manda-karṇa. It can be shown that the iterated manda-karṇa is actually given (in the limit) by OPin Figure 5, where the point P is obtained as follows.¹¹ Consider a point O'' at a distance of r_0 from O along the direction of mandocca OU and draw O''P₁ so that it meets the concentric at Q. Then produce OQ to meet the extension of P_0P_1 at P.

Though there is no need to evaluate the karna to compute $\Delta\theta$, the mandakarna K is needed in other calculations. For instance, in eclipse computation we need to find the true distance of the Sun or Moon. Similarly, if we have to find the latitude of the planet, we need the heliocentric distance of the planet, which is given by avisista-manda-karna (iterated hypotenuse). Needless to say that it would have been an arduous task to determine avisista-manda-karna by iterative procedure in those days, given that fact that it involves repeated computation of squares and square roots.

To circumvent this difficult exercise of performing an iterative process, Mādhava, by carefully analysing the geometry of the problem came up with a brilliant method for finding the *aviśiṣṭa-manda-karṇa*. This method due to Mādhava involves only one step and hence is usually referred to as *sakṛt-karma* as opposed to the *asakṛt-karma* prescribed by Bhāskara.

4.2.3 Mādhava's method for obtaining the *aviśiṣṭa-manda-karṇa* through *sakṛt-karma*

The verses commencing with *vistrti*, which as mentioned earlier forms one of the $m\bar{u}la$ verses of the manuscript, succinctly presents the *sakrt-karma* method of finding the *avisista-manda-karna* given by Mādhava.

¹¹See for instance, the discussion in {MB 1960}, pp. 111–9.

विस्तृतिदलदोःफलकृतिवियुतिपदं कोटिफलविहीनयुतम्। केन्द्रे मृगकर्किगते स खलु विपर्ययकृतो भवेत् कर्णः॥ तेन हृता त्रिज्याकृतिः अयत्नविहितोऽविशेषकर्णः स्यात् ।

The square of the *dohphala* is subtracted from the square of the $trijy\bar{a}$ and its square root is taken. The *koțiphala* is added to or subtracted from this depending upon whether the *kendra* (anomaly) is within six signs beginning from *Karki* (Cancer) or *Mrga* (Capricorn). This gives the *viparyayakarṇa*. The square of the *trijyā* divided by this *viparyayakarṇa* is the *aviśeṣakarṇa* (iterated hypotenuse) obtained without any effort [of iteration].

As described in the above verse, Mādhava's method involves finding a new quantity called the *viparyayakarna* or *viparītakarna*. The term *viparītakarna* literally means 'inverse hypotenuse', and is nothing but the radius of the *kakṣyāvṛtta* when the measure of *mandakarna* is taken to be equal to the *trijyā*, R.

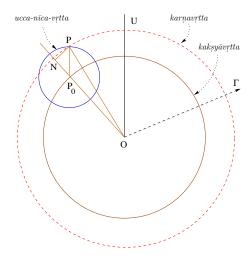


Figure 6: Determination of the *viparītakarņa* when the *kendra* is in the first quadrant.

The rationale behind the formula for *viparītakarņa*, presented in this manuscript under study, is also explained in the celebrated Malayalam text *Yuktibhāṣā*,¹² and can be understood with the help of Figures 5 and 6. In these figures P_0 and P represent the mean and the true planet respectively. N denotes the foot of a perpendicular drawn from the true planet P to the line joining the centre of the

 $^{^{12}{\}rm See}$ {GYB 2008}, pp. 484–6, 635–40.

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circle and the mean planet. NP is equal to the *dohphala*. Let the radius of the *karnavitta* OP be set equal to the *trijyā* R. Then the radius of the *uccanīcavitta* P_0P is r_0 , as it is in the measure of the *karnavitta*. In this measure, the radius of the *kakṣyāvitta* $OP_0 = R_v$, the *viparītakarṇa*, and is given by

$$R_{v} = ON \pm P_{0}N$$

= $\sqrt{R^{2} - (r_{0}\sin(\theta_{0} - \theta_{m}))^{2}} \pm |r_{0}\cos(\theta_{0} - \theta_{m})|.$ (10)

Mādhava, besides besides giving an expression for R_v in terms of the mean anamoly $(\theta_0 - \theta_m)$, as in (10), also seems to have provided an an alternative expression for R_v in terms of the true anamoly $(\theta_{ms} - \theta_m)$, as follows:

$$R_v = \sqrt{R^2 + r_0^2 - 2r_0 R \cos(\theta_{ms} - \theta_m)}$$
(11)

This is clear from the triangle OP_0P , where $OP_0 = R_v$, OP = R and $P_0PO = \theta_{ms} - \theta_m$.

In Figure 5, Q is a point where $O''P_1$ meets the concentric. OQ is produced to meet the extension of P_0P_1 at P. Let T be the point on OP_0 such that QTis parallel to P_0P_1 . Then it can be shown that $OT = R_v$ is the *viparītakarņa*. Since triangles OQT and OPP_0 are similar, we have

$$\frac{OP}{OP_0} = \frac{OQ}{OT} = \frac{R}{R_v}$$

or, $OP = K = \frac{R^2}{R_v}.$ (12)

Thus we have obtained an expression for the *avisista-manda-karṇa* in terms of the *trijyā* and the *viparītakarṇa*. Once we find R_v either using (10) or (11), the *avisista-manda-karṇa* K can be calculated using (12) and thereby avoid the iterative process.

4.3 Explanation of the verse 'arkendvoh'

The verse commencing with arkendvoh—the second of the two $m\bar{u}la$ verses commented in the manuscript—outlines the procedure for obtaining the mean positions (madhyama) of the Sun and the Moon from their true positions (mandasphuța). This verse, originally due to Mādhava and quoted by later Kerala astronomers, runs as follows: अर्केन्द्रोः स्फुटतो मृदूचरहितात् दोःकोटिजाते फले नीत्वा कर्किमृगादितो विनिमयेनानीय कर्णं सकृत् । त्रिज्यादोःफलघाततः श्रुतिहृतं चापीकृतं तत् स्फुटे केन्द्रे मेषतुलादिगे धनमृणं तन्मध्यसंसिद्धये ॥ ४१ ॥

Subtracting the longitude of their own mandoccas from the true positions of the Sun and the Moon, obtain their dohphala and koțiphala. Find the sakrt-karṇa (one-step hypotenuse) once by interchanging the sign [in the cosine term] depending upon whether the kendra is within the six signs beginning with Karki or Mrga. Multiplying the dohphala and trijyā, and dividing this product by the karṇa [here referred to as śruti], the arc of the result is applied to the true planet to obtain the mean planet. This arc has to be applied positively and negatively depending upon whether the kendra lies within the six signs beginning with Meṣa or Tulā respectively.

The term $mrd\bar{u}cca$ appearing in the first line of the verse is a synonym for mandocca. The sphuta that is referred to here should be understood as manda-sphuta, since it has been explicitly mentioned in the verse that the procedure outlined here is applicable only for the Sun and the Moon.¹³ If θ_{ms} and θ_m represent the longitudes of the manda-sphuta and mandocca (of the Sun or the Moon), then their sphuta-dohphala and sphuta-kotiphala are given by

$$dohphala = r_0 \sin(\theta_{ms} - \theta_m)$$

$$kotiphala = r_0 \cos(\theta_{ms} - \theta_m), \qquad (13)$$

where r_0 represents the radius of the mean epicycle whose values are provided in the text. With these *doḥ* and *koṭiphalas*, the *sakṛt-karṇa* (expression similar to (2), with opposite sign in *koṭiphala*), may be written as

$$karna = [(R - r_0 \cos(\theta_{ms} - \theta_m))^2 + (r_0 \sin(\theta_{ms} - \theta_m))^2]^{\frac{1}{2}}$$
(14)

It can be easily seen that the above expression is the same as the expression for the *viparīta-karṇa* R_v given by (11). The first half of the above verse essentially prescribes to obtain the expression for R_v . We now proceed to explain the second half of it with the help of Figure 7.

In Figure 7, O is the observer and P_0 is the mean planet (the Sun or the Moon). The point P represents their true position. The distance $P_0P = OO'$

 $^{^{13}}$ The reason as to why the domain of applicability is restricted only to the Sun and the Moon, is explained by Nīlakaṇṭha in his $\bar{A}ryabhaṭ\bar{i}ya-bh\bar{a}sya$. We discuss this in the next section.

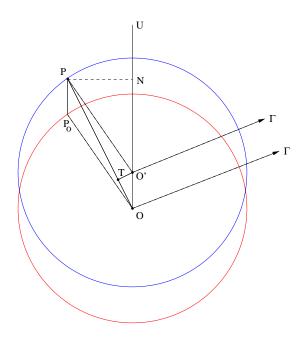


Figure 7: Obtaining the *madhyama* (the mean position) from the *sphuța* (the true position).

represents the radius of the actual variable epicyle that we denote as r. The angles $P_0 \hat{O}P = OPO' = (\theta_{ms} - \theta_m)$. Considering the triangle OO'P, we draw a perpendicular from O' that intersects OP at T. Now, in the triangle O'PT,

$$O'T = O'P\sin(O'\hat{P}T)$$

= $O'P\sin(P\hat{O}P_0)$
= $R\sin(\theta_0 - \theta_{ms}).$ (15)

Also
$$O'T = r\sin(\theta_{ms} - \theta_m).$$
 (16)

Equating the above two expressions for O'T,

$$R\sin(\theta_0 - \theta_{ms}) = r\sin(\theta_{ms} - \theta_m)$$

or
$$R\sin(\theta_0 - \theta_{ms}) = r_0\sin(\theta_{ms} - \theta_m)\frac{R}{R_v},$$
 (17)

where we have used (2.135) and (2.153). Hence,

$$\theta_0 - \theta_{ms} = (R\sin)^{-1} \left[r_0 \sin(\theta_{ms} - \theta_m) \frac{R}{R_v} \right].$$
(18)

Thus the mean planet θ_0 can be obtained by adding the above difference to the true planet θ . $\theta_0 - \theta_{ms}$ is positive when the *kendra* (anomaly) $\theta_{ms} - \theta_m$

is within the six signs beginning with *Mesa*, viz., $0^{\circ} \leq \theta_{ms} - \theta_m \leq 180^{\circ}$, and negative when the *kendra* is within the six signs beginning with *Tulā*, viz., $180^{\circ} \leq \theta_{ms} - \theta_m \leq 360^{\circ}$.

Having briefly explained the content of the verse, we now proceed to discuss it in greater detail in the light of the edifying commentary authored by Nīlakanṭha on it in his $\bar{A}ryabhaṭiya-bhāṣya$. As the commentary is quite elaborate, runs to more than two pages, we only present excerpts from it. The excerpts have been chosen primarily to highlight some of special features pointed out be Nīlakanṭha, particularly with respect to the distinction that must be maintained in applying the procedure, outlined by the above verse, for obtaining the *manda-sphuṭa* from the śighra-sphuṭa of the planets.

4.3.1 Obtaining the manda-sphuta from the śighra-sphuta

Since the geometrical construction involved in the manda as well as the \hat{sighra} process is essentially the same—both just involve a deferent circle and an epicyle, though the significance of them widely varies in the two processes— Nīlakaṇṭha, in his $\bar{A}ryabhațīya-bhāṣya$, outlines the procedure for obtaining the manda-sphuṭa from the $\hat{sighra}-sphuṭa$.

अत एव शीघ्रस्फुटेन स्फुटमध्यमानयने शीघ्रस्फुटतदुचविवरभुजां स्वपरिधिहतां चक्रांशैः अशीत्या वा हृत्वा लब्यं दोःफलं केवलमेव चापीकृत्य शीघ्रस्फुटे व्यत्ययेन संस्कुर्यात्। तदा स्फुटमध्यमं स्यात्।...

अत उत्तं - 'मृद्चरहितादि'ति।

It is for this reason, it has been prescribed that the arc corresponding to the *dohphala*—obtained from the Rsine of the difference between the $\hat{sighra-sphuta}$ and $\hat{sighrocca}$ multiplied by its own circumference (*sva-paridhihatām*¹⁴) and divided by either 360 or 80—should be applied to the $\hat{sighra-sphuta}$ inversely, in order to the *manda-sphuta* from the $\hat{sighra-sphuta}$.

We may explain the content of the above passage with the help of Figure 5. In this figure, O is the observer, S is the Sun and P any of the three exterior planets Mars, Jupiter or Saturn. $OS = r_s$ represents the radius of the $s\bar{s}ghra$ epicycle

 $^{^{14}}$ Since the discussion is on the $\dot{s\bar{\imath}g}hra$ process, the word 'sva' refers to the $\dot{s\bar{\imath}g}hra$.

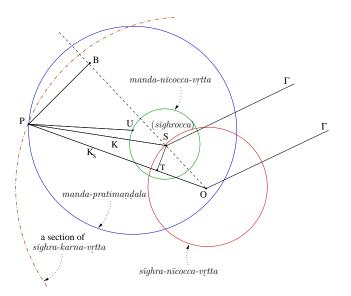


Figure 8: Obtaining the manda-sphuta from the śīghra-sphuta.

and SP = K the *aviśiṣṭa-manda-karṇa*. The angles $\Gamma \hat{S}P = \theta_{ms}$ and $\Gamma \hat{O}P = \theta$ represent the *manda-sphuṭa* and *śīghra-sphuṭa* respectively. θ_s represents the longitude of *śīghrocca* which is the same as the longitude of the mean Sun that is known. The objective is to find the θ_{ms} from θ .

Considering the triangle OPS, and applying the sine formula we have

$$\frac{r_s}{\sin(\theta_{ms} - \theta)} = \frac{K_s}{\sin(\theta_{ms} - \theta_s)} = \frac{K}{\sin(\theta - \theta_s)}.$$
(19)

Hence,

$$\sin(\theta_{ms} - \theta) = \frac{r_s}{K} \sin(\theta - \theta_s)$$

or, $(\theta_{ms} - \theta) = \sin^{-1} \left(\frac{r_s}{K} \sin(\theta - \theta_s)\right)$
or, $\theta_{ms} = \theta + \sin^{-1} \left(\frac{r_s}{K} \sin(\theta - \theta_s)\right).$ (20)

Thus the manda-sphuta can be obtained from the $s\bar{i}ghra-sphuta$ using the above relation.

T16

4.3.2 Applicability of the procedure outlined in the verse *arkendvo*^h, to other planets

Though by conceding the analogy between the geoemtrical constructions involved in the manda and the $\hat{sig}hra$ processes, though Nīlakaṇṭha prescribed a procedure for getting the manda-sphuṭa from the $\hat{sig}hra-sphuṭa$, he clearly maintains the distinction between the two. In fact, to avoid any confusion he explicitly points out that the sakrt-karma procedure given for the Sun and the Moon cannot be simply generalized to find the madhyama from the sphuṭa in the case of the planets Mars, Jupiter and so on. He says:¹⁵

अर्केन्द्रग्रहणं भौमादीनां निवृत्त्यर्थम्। तेषां दोर्ज्यावशात् परिधिभेदात् ...

Specific mention of the Sun and the Moon is to desist the entry of Mars, etc. For them, since the circumference changes as per $dorjy\bar{a}$, ...

The justification presented by Nīlakantha as to why the procedure outlined by Mādhava for obtaining the the mean position from the true position using *sakrt-karma*, is not applicable to the other five planets is quite involved and requires an elaborate explanation. As this does not fall under the scope of the present paper, we skip further discussion on this topic and move on to the next section.

5 The Edited text and translation

The edited version of the text presented below, is identical with the manuscript but for the insertion of numbers before the beginning of the paragraphs. These numbers have been inserted for the purpose of easily identifying the original text with the explanatory notes that has been prepared in four sections.¹⁶ The sectioning too has been done in manner that seemed most appropriate and convenient for our discussion.

(1) विस्तृतीत्यादि। अर्केन्द्रोरित्यादि। अनयोः श्लोकद्रययोर्युत्तिप्रदर्शनाय त्रिज्या-तुल्यव्यासार्थं कक्ष्यामण्डलमालिख्य मातुपितुरेखां कृत्वा तत्परिधौ रविमध्यमा-

 $^{^{15}\{ {\}rm ABB \ 1931} \},$ p. 50; This passage is quoted in a different context in $\{ {\rm LB \ 1979} \}$ as well.

¹⁶Para (1) forms the first section, Para (2) the second, Para (3)–(5) the third and Para (6) the last.

धिष्ठितप्रदेशं केन्द्रं कृत्वा अविशिष्ठान्त्यफलत्रिज्यातुल्येन सूत्रेण उच्चनीचवृत्त-मालिख्य तत्परिधौ मन्दोचदिशि स्फुटग्रहमालिख्य ग्रहाधिष्ठितप्रदेशात कक्ष्या-मण्डलकेन्द्रप्रापिणं सूत्रं कुर्यात्। कक्ष्यामण्डलकेन्द्रात् स्ववृत्तगतोच्रप्रापिणं सूत्रं च कुर्यात्। पुनः कक्ष्यामण्डलकेन्द्रमेव केन्द्रं कृत्वा अविशिष्टान्त्यफलसूत्रेण द्वितीयमप्युचनीचवृत्तमालिखेत्। पुनः पूर्वलिखितोच्चनीचपरिधेः¹⁷ कर्णवृत्तपरिधेश्च सम्पातद्वयप्रापिणं सूत्रमालिखेत्। तदर्धं मन्दोचोनमध्यमाल्लखं¹⁸ अविशिष्टभुजा-फलम्। एतत्सूत्रमध्यात् मध्यमग्रहप्रापिसूत्रं अविशिष्टमध्यमकोटिफलम्। पुनर्द्वि-तीयोच्चन्नीचवृत्तस्य कक्ष्यामण्डलस्य च परिधिसम्पातद्वयप्रापिणं सूत्रं कुर्यात्। तदर्धं मन्दोचोनात् अर्कस्फुटाल्लब्यं अविशिष्टभुजाफलम्। तन्मध्यात् स्फुटग्रहप्रापि सूत्रमविशिष्टस्फुटकोटिफलम्। स्फुटभुजाकोटिफलाभ्यां कर्णानयनं त्रिज्यावर्गात् भुजाफलवर्गं विशोध्य मूले¹⁹ कोटिफलं मृगकर्क्यादितः स्वर्णं कृत्वा साध्यम्। तस्य कर्णस्य पुनरविशेषणं पूर्ववदेव।

(2) एवं साधनचतुष्टये लिखिते तैर्विषमचतुरश्रं क्षेत्रं भवति। तत्र मध्यमोद्भवयोः भुजाकोटिफलयोः स्फुटोद्भवयोश्च साधारणं कर्णो मध्यमग्रहात् स्फुटग्रहप्रापि-सूत्रम्। एवं क्षेत्रे लिखिते कक्ष्यामण्डल स्फुटग्रहप्रापिकर्णसूत्रात्²⁰ त्रिज्याया आनयनं²¹ निरूप्यम्। तदाथा कर्णवर्गात् मध्यमदोःफलवर्गं विशोध्य यन्मूलं तस्मिन् मध्यमकोटिफलं मृगकर्क्यादावृणं धनं च कृत्वा ति्रज्या²² लभ्यते ²³ कर्णे स्फुटकोटिफलं मृगकर्क्यादावृणं धनं च कृत्वा तद्वर्गे²⁴ स्फुटदोःफलवर्गं क्षिप्त्वा मूलं च त्रिज्या भवति। एवं द्विविधं त्रिज्यानयनमवगम्य पुनर्विपरीतकर्णा-नयनं निरूप्यते।

(3) विपरीतकर्णस्तु क इति चेत्, कर्णे त्रिज्यातुल्ये सति कियती त्रिज्या इति लब्धा त्रिज्या विपरीतकर्णः। उक्तप्रकारेण तदानयनाय अवशिष्टस्य वृत्तगते मध्यमकोटिभुजाफले स्फुटकोटिभुजाफले अविशिष्टकर्णं त्रिज्यां च त्रिज्यया हत्वा कर्णेन हरेत्। तदानीं कर्णस्त्रिज्यातुल्यो भवति। मध्यमकोटिभुजाफले प्रथमं मन्दोद्योनमध्यमाल्लब्ययोर्बाहुकोटिज्ययोः परिधिगुणनाशीतिहरणे कृत्वा लब्याभ्यां भुजाफलकोटिफलाभ्यां तुल्ये च भवतः। त्रिज्या, त्रिज्यावर्गं²⁵ अविशिष्टकर्णेन

¹⁹मूलं

²³Here the sentence "कर्णे स्फुटकोटिफलं मृगकर्क्यादावृणं धनं च कृत्वा वर्गं त्रिज्या लम्यते।" appears twice in the manuscript.

²⁴तद्वर्गगे

25 त्रिज्यालगार्गत्रिज्य

¹⁷नीच परिसं

¹⁸मन्दोचोनमा लब्धं

²⁰ प्रापिणं कर्णसूत्रात्

²¹आनयनयन्

²²कृत्वा वर्गं त्रिज्या

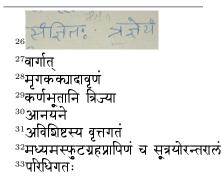
हृत्वा लब्धेन तुल्या भवति। तद्विपरीतकर्णसंज्ञितोऽत्र ज्ञेयम् ²⁶ तदर्थं त्रिज्या-भूतस्य कर्णस्य वर्गात्²⁷ मध्यमभुजाफलवर्गमपहाय यन्मूलं तस्मिन् मध्यम-कोटिफलं मृगकर्क्यादावृणं²⁸ धनं कुर्यात्। ततो विपरीतकर्णभूता त्रिज्या²⁹ लभ्यते। अत उक्तं 'विस्तृतिदले'त्यादिना। त्रिज्याभूते कर्णे सकृत्कर्मणा स्फुटतो लब्धं कोटिफलं मृगकर्कितः ऋणं धनं च विधाय तद्वर्गे सकृत्कर्मलब्धस्य स्फुटभुजाफलस्य वर्गं क्षिप्त्वा मूलं च विपरीतकर्णभूता त्रिज्या भवति। अत उक्तं – 'अर्केन्द्रोः स्फटत' इत्यादिना।

(4) एवं विपरीतकर्णे ज्ञाते अविशेषकर्णानयनाय विपरीतकर्णमूतायां त्रिज्यायां त्रिज्यातुल्यायां जातायां त्रिज्यामूतः कर्णः कियानिति त्रैराशिकम्। तत्र फलस्य इच्छायांश्च त्रिज्या तुल्यत्वात् त्रिज्यावर्गो विपरीतकर्णेन भाज्यम्। लब्धं अविशेष-कर्णश्च भवति। अत उक्तं – तेन ह्वता त्रिज्याकृतिरयत्नविहितोऽविशेषकर्णः स्यात्। – इति।

(5) एवं मध्यमतः स्फुटतश्च सकृत्कर्मणा अविशेषकर्णलब्धिश्च भवति। त्रिज्या-विशेषकर्णाभ्यां त्रैराशिके कर्तव्ये सर्वत्र त्रिज्यास्थाने विपरीतकर्णः, अविशेष-कर्णस्थाने त्रिज्यां च कृत्वा कर्म करणीयं भवति। तत्र स्फुटतो मध्यमानयनं³⁰ करिष्यते।

(6) एवं कर्णानयने ज्ञाते स्फुटतो मध्यमानयनं निरूप्यते – पूर्वलिखितमविशिष्ट-वृत्तगतं³¹ स्फुटभुजाफलं कक्ष्यामण्डलकेन्द्रात् मध्यमस्फुटग्रहप्रापिसूत्रयोरन्त-रालं³²कक्ष्यामण्डलपरिधिगतम्³³। अतः अविशिष्टस्फुटभुजाफलमेव मध्यम-स्फुटान्तरम्। तज्ज्ञानाय सकृत्कर्मलब्धं स्फुटभुजाफलमविशेषकर्णेन हत्वा त्रिज्यया हर्तव्यम्। अथवा त्रिज्यया हत्वा विपरीतकर्णेन विभज्यापि लम्यते। तब्यचापं मेषादौ मध्यमसूत्रयोः मध्यमसूत्रस्याग्रगतत्वात् स्फुटे धनं कार्यम्। तुलादौ पृष्ठगतत्वात् ऋणं च कार्यम्। अतः – 'त्रिज्यादोःफलघाततः श्रुतिहृतं चापीकृतं तत्स्फुटे केन्द्रे मेषतुलादिगे धनमुणं तन्मध्यसंसिद्धये' इत्युक्तम्।

(1) [The verse] that commences with *vistrti*. [The verse] that commences with *arkendvoh*. In order to present the rationale behind



the [content] of these two verses, having drawn the deferent circle (kaksyā-mandala) with radius equal to the trijyā, mark the eastwest line $(m\bar{a}t\bar{r}-pit\bar{r}rekh\bar{a})$ in it. Then, having marked the position of the mean Sun on its circumference, draw an epicycle with that as centre and with the radius equal to *aviśistāntyaphala*. Marking the position of the true planet on its circumference along the direction of the apogee (mandocca), draw a line from the location of the planet to the center of the deferent circle. Also draw a line, that reaches the apogee corresponding to that circle from the center of the kaksyā-mandala. Again, with the center of the deferent circle as the center, draw a second epicycle with radius equal to the aviśistāntya-phala. Then, draw a line connecting the points of intersection of the circumference of the epicycle that was drawn earlier and the circumference of the karnavrtta. Half of that [line] is aviśista-bhujāphala obtained from the mean Sun diminished by the apogee. The line reaching the mean planet from the center of this line is the *aviśistakotiphala* of the mean planet. Then, draw a line joining the points of intersection of the circumference of the second epicycle and the circumference of the kaksyāmandala. Half of that [line is] avisista-bhujāphala obtained from the true Sun diminished by the apogee. The line reaching the true planet from the middle of this line [is the] aviśistakotiphala of the true [planet]. The karna is obtained from the true $bhuj\bar{a}$ and kotiphalas by first finding the square root of the difference of the squares of the bhujāphala and the $trijy\bar{a}$ and applying the *kotiphala* to that positively or negatively depending upon [whether the kendra is] Mrgādi or Karkyādi respectively. The iteration of that karna is to be done as before.

(2) Thus, the four means $(s\bar{a}dhanacatustaya)$ [namely, the bhujā and the kotiphalas corresponding to the mean and the true planets] when represented, form a quadrilateral of unequal sides (visamacaturasta). In that, the line joining the mean and the true planet, forms the common hypotenuse of the bhujā and the kotiphalas corresponding to the mean and the true [kendras]. Thus, having drawn the figure, now the procedure for obtaining the trijyā from the karṇa, which joins the centre of the kakṣyāmaṇḍala and the true planet, is to be explained. This is as follows: By applying the madhyama-kotiphala negatively or positively—depending upon whether the kendra is Mrgādi or Karkyādi respectively—to the the square root of the difference of the the squares of the karṇa and the madhyama-doḥphala, trijyā is obtained. [Similarly] having applied the true kotiphala to the karṇa negatively or positively depending upon whether the kendra is $Mrg\bar{a}di$ or $Karky\bar{a}di$ respectively, the square of that has to be taken. By adding the square of the *sphuta-dohphala* to the square of that, and taking the square root, $trijy\bar{a}$ is obtained. Thus, having understood the two ways of obtaining the $trijy\bar{a}$, we now proceed to explain the procedure for obtaining the inverse hypotenuse.

(3) [If you ask] what *viparītakarna* is, [we say], when the *karna* is taken to be equal to the measure of the $trijy\bar{a}$, then the value of $trijy\bar{a}$ obtained in that measure is the $vipar\bar{i}takarna$. To obtain that along the lines described above multiply the $trijy\bar{a}$ with itself [and] divide by the karna. Then, karna will be equal to the $trijy\bar{a}$. The $bhuj\bar{a}$ and the kotiphalas [thus obtained] will be equal to the $b\bar{a}hu$ and the $kotijy\bar{a}s$ obtained earlier by subtracting the apogee from the mean planet, when multiplied by the circumference of the epicycle and divided by eighty. It is to be known here, that the quantity obtained by dividing the square of the $trijy\bar{a}$ by the aviśistakarna, is the viparītakarna. In order to obtain that, we first find the square root of the difference of the squares of karna taken in the measure of the $trijy\bar{a}$ and the square of madhyamabhujāphala. To the square root of that, the madhyama-kotiphala is applied. negatively or positively depending upon whether the kendra is Mrqādi or Karkyādi. Now, the viparīta-karna in the measure of $trijy\bar{a}$ is obtained. It is therefore said "vistrtidala" etc. To the karna which is in the measure of $trijy\bar{a}$, apply the sphuta-kotiphala obtained through *sakrt-karma* negatively or positively depending upon whether the kendra is Mrqādi or Karkyādi respectively. To the square of that, add the square of the *sphuta-bhujāphala* [again] obtained through *sakrt-karma*, and find its square root. The result will be the *viparītakarna*, in the measure of the *trijyā*. It is therefore said - "arkendvoh sphutatah".

(4) Thus having known the *viparīta-karņa*, the rule of three that is to be employed for obtaining the iterated hypotenuse is : If the *viparīta-karņa* were to be taken to be equal to the measure of the *trijyā*, then what would be the value of the *karṇa* which was previously taken to be equal to the *trijyā*. Since here, the *pramānaphala* as well as *icchā* are equal to the *trijyā*, the square of the *trijyā* has to be divided by the *viparīta-karṇa*. The resultant is the iterated hypotenuse. It is therefore said – 'The square of the *trijyā* divided by that [*viparyayakarṇa*] is the *aviśeṣakarṇa* (iterated hypotenuse) obtained without any effort [of doing iteration]'.

(5) Thus either from the mean or the true the iterated hypotenuse is

obtained through the *sakṛt-karma*. Since the rule of three should be applied using the $trijy\bar{a}$ and the iterated hypotenuse, all the operations have to be executed by replacing $trijy\bar{a}$ with the $vipar\bar{\imath}ta$ -karṇa and the aviśeṣakarṇa with the $trijy\bar{a}$. Now, the mean [planet] from the true will be obtained.

(6) Thus having known how to find the karna, the procedure for obtaining the mean from the true is now explained. The sphuta*bhujāphala* which lies inside the iterated epicycle drawn earlier, which is essentially equal to the distance of seperation between the lines drawn from the centre of the $kaksy\bar{a}$ -mandala, is [also] inside the circumference of the kaksyā-mandala. Hence, the aviśista $sphuta-bhuj\bar{a}phala$ is the same as difference between the mean and the true [positions of the planet]. In order to obtain that, the *sphuta*bhujāphala obtained through the sakrt-karma should be multiplied by the iterated hypotenus and divided by the $trijy\bar{a}$. The same result may also be obtained by multipling by the $trijy\bar{a}$ and dividing by the viparitakarna. If the kendra is Mesādi, then the resulting arc should be applied positively to the true planet, since of the two lines joining the mean and the true [from the centre of the kaksyāmandala], the one joining the mean leads the other. If the kendra is $Tul\bar{a}di$, then the arc should be applied negatively since the line joining the mean planet lags behind the other. It is therefore said that – 'Multiplying the dohphala and $trijy\bar{a}$, and dividing this product by the karna [here referred to as *śruti*], the arc of the result is applied to the true planet to obtain the mean planet. This arc has to be applied positively and negatively depending upon whether the kendra lies within the six signs beginning with Mesa or $Tul\bar{a}$ respectively.'

6 Explanatory notes

Transcending the confines of immediate utility of merely explaining the content of the verses, the commentary in the manuscript gradually develops the background material that would enable the reader to have a fuller appreciation of the content of the $m\bar{u}la$. It also attempts to present the rationale behind the procedures outlined in the verses. In what follows, we present the content of the manuscript using modern mathematical notations. For the purpose of convenience we have divided it into four sections.

6.1 Definition of *sādhanacatustaya*

The manuscript commences with detailed description of how the geometrical figure needs to be constructed with which the two verses to be commented upon can be understood. The geometrical figure described therein may be depicted as indicated in Figure 9.

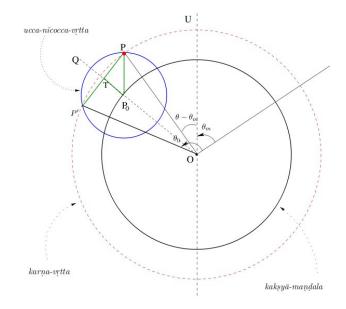


Figure 9: The geometrical construction described in the manuscript as a tool to understand the $m\bar{u}la$ verses.

Here O represents the centre of the deferent circle known as kaksyamandala, whose radius is R. The mean planet P_0 is located on this circle, whose longitude is denoted as θ_0 . The circle centered around P_0 and with radius r_0 is called *uccanīca-vrtta* (epicycle). A line parallel to the direction of apogee (OU) drawn from the P_0 meets the epicycle at P. This gives the position of the manda-sphuta (the manda corrected planet). Now, the circle centered at O and having a radius OP = K, known as karna-vrtta is drawn. This intersects the epicycle at two points, namely P and P'. The line OQ, which is an extension of the line joining the centre of the deferent circle and the mean planet bisects the line PP' at T. Now considering the $\triangle PP_0T$,

 $P\hat{P}_0T = \theta_0 - \theta_m,$

where θ_0 and θ_m are longitudes of mean planet and apogee respectively. In this triangle which is a right-angled at the vertex T the hypotenuse P_0 represents the *aviśiṣṭāntyaphala* (r). Thus, we have

$$PT = r\sin(\theta_0 - \theta_m) \tag{21}$$

$$P_0T = r\cos(\theta_0 - \theta_m) \tag{22}$$

The quantities PT and P_0T , defined in the above equations are referred to as avisista-madhyama-bhujaphala and avisista-madhyama-kotiphala respectively.

A section of Figure 9 is blown up and depicted in Figure 10. Here $OP_0 = R$ represents the *trijyā* and OP = K denotes the *manda-karṇa*. The foot of perpendicular drawn from P_0 to OP, intersects the latter at B. In the $\triangle PP_0B$, $P_0\hat{P}B = \theta_{ms} - \theta_m$. Hence,

$$P_0 B = r \sin(\theta_{ms} - \theta_m) \tag{23}$$

$$PB = r\cos(\theta_{ms} - \theta_m). \tag{24}$$

These two quantities P_0B and PB are known as *sphutabhuja* and *sphutakoti* respectively. It may be noted that the two triangles constitute a quadrilateral of unequal sides. It is the four sides of this quadrilateral—representing the *bhujā* and *koți* of the *madhyama* and *sphuta-kendras*—that are referred as $s\bar{a}dhanacatuśtaya$. Literally the term *sadhanacatuśtaya* means "a group consisting of four-means".

Having defined sadhanacatustaya, the text presents a formula for the karna OP in terms of the sphutabhujā and kotiphalas. In the $\triangle OP_0B$,

$$OB = \sqrt{OP_0^2 - P_0 B^2}$$
$$= \sqrt{R^2 - (r\sin(\theta_{ms} - \theta_m))^2}$$

Also, OB = OP - BP. Hence,

$$OP = \sqrt{R^2 - (r\sin(\theta_{ms} - \theta_m))^2} + BP$$
(25)

$$K = \sqrt{R^2 - (r\sin(\theta_{ms} - \theta_m))^2 + r\cos(\theta_{ms} - \theta_m)}.$$
 (26)

Thus manda-karna can be obtained in terms of trijyā and sphuta-kendra.

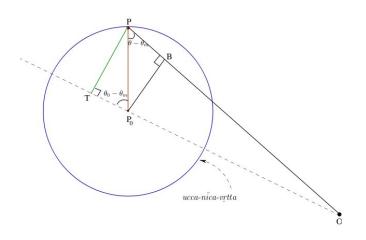


Figure 10: The sādhana-catustaya employed in finding the manda-karņa.

6.2 Obtaining $trijy\bar{a}$ from karna

Having defined a relation to obtain the karna K, in terms of the trijyā R, the commentary presents two methods by which one can obtain $trijy\bar{a}$ in terms of the karna.

Method 1: Here $trijy\bar{a}$ is obtained in terms of the karna and the madhymabhuj \bar{a} -kotiphalas. The expression given can be understood with the help of Figure 10. Considering the $\triangle OTP$, we have

$$OT^{2} = OP^{2} - PT^{2}$$

or $(OP_{0} + P_{0}T)^{2} = K^{2} - (r\sin(\theta_{0} - \theta_{m}))^{2}$ (27)

Hence,

$$OP_{0} = \sqrt{K^{2} - (r\sin(\theta_{0} - \theta_{m}))^{2}} - P_{0}T$$

or
$$R = \sqrt{K^{2} - (r\sin(\theta_{0} - \theta_{m}))^{2}} - r\cos(\theta_{0} - \theta_{m}).$$
 (28)

Method 2: In this method R is found in terms of K and the *sphuțabhujā-koțiphalas*. In the $\triangle OP_0B$ (see Figure 10),

$$OP_0^2 = OB^2 + BP_0^2$$

or $R^2 = (OP - BP)^2 + (r\sin(\theta - \theta_m))^2$
or $R = \sqrt{(K - r\cos(\theta - \theta_m))^2 + (r\sin(\theta - \theta_m))^2}.$ (29)

6.3 Expression for *viparītakarņa* and its application

The third paragraph of the $m\bar{u}la$ starts with the definition of $vipar\bar{\iota}takarna$ and then proceeds to give two different expressions for the same—one in terms of the $trijy\bar{a}$ and the manda-kendra and the other in terms of the $trijy\bar{a}$ and the sphuta-kendra. The definition given here may be stated as follows: If the measure of the karna were to be taken to be equal to the $trijy\bar{a}$ in a certian scale, then whatever that turns out to be the magnitude of $trijy\bar{a}$ in the same scale is defined as $vipar\bar{\iota}ta$ -karna. Symbolically this may be represented as a problem of rule of three:

$$K : R$$

$$R : R_v(?)$$
or
$$R_v = \left(\frac{R}{K}\right)R$$
(30)

The two expressions for R_v given here are the same as the equations (10) and (11) discussed in section 4.2.3. As a detailed derivation of the two equations are presented there itself, we do not repeat the same here. It would suffice to recount that one of the main purposes for introducing this mathematical device R_v is to find K without having to resort to *aviśeṣa-karma*. Para (4) essentially states that *aviśiṣta-manda-karṇa* or *aviśeṣa-karṇa* K can be expressed in terms of *trijyā* and *viparīta-karṇa* as:

$$aviśeṣakarṇa = \frac{trijy\bar{a}^2}{viparītakarṇa}$$

or $K = \frac{R^2}{R_v}.$ (31)

In Para (5) it is prescribed that the $trijy\bar{a}$ and avises a karna can be replaced by the $vipar\bar{\iota}ta$ -karna and $trijy\bar{a}$ respectively in all the operations that are to be carried out based on rule of three. That is,

The above prescription can be better appreciated with the help of relations that can be derived from the triangles OP_0P and OTQ in the Figure 5. Considering these two triangles and applying the sine formula we have

$$\frac{R}{\sin(\theta_{ms} - \theta_m)} = \frac{K}{\sin(\theta_0 - \theta_m)} = \frac{r}{\sin(\theta_0 - \theta_{ms})}$$
(33)

$$\frac{R_v}{\operatorname{in}(\theta_{ms} - \theta_m)} = \frac{R}{\operatorname{sin}(\theta_0 - \theta_m)} = \frac{r_0}{\operatorname{sin}(\theta_0 - \theta_{ms})}.$$
(34)

From (33) the following equations may be obtained.

$$\sin(\theta_0 - \theta_m) = \frac{K}{R}\sin(\theta_{ms} - \theta_m)$$
(35*a*)

$$\sin(\theta_0 - \theta_{ms}) = \frac{r}{K}\sin(\theta_0 - \theta_m)$$
(35b)

$$\sin(\theta_0 - \theta_{ms}) = \frac{r}{R}\sin(\theta_{ms} - \theta_m). \tag{35c}$$

Similary from (34), we have

$$\sin(\theta_0 - \theta_m) = \frac{R}{R_v} \sin(\theta_{ms} - \theta_m)$$
(36*a*)

$$\sin(\theta_0 - \theta_{ms}) = \frac{r_0}{R} \sin(\theta_0 - \theta_m)$$
(36b)

$$\sin(\theta_0 - \theta_{ms}) = \frac{r_0}{R_v} \sin(\theta_{ms} - \theta_m)$$
(36c)

It may be observed that LHS, in the pair of three sets of equations presented above, is one and the same. So too is the argument of the sine function in RHS. This only forces us to conclude that multipying factors—in the form of ratios of two quantities—that appear in the RHS must also be the same though they look apparently different. Thus we have

$$\frac{K}{R} = \frac{R}{R_v}; \qquad \frac{r}{K} = \frac{r_0}{R}; \qquad \text{and} \qquad \frac{r}{R} = \frac{r_0}{R_v}, \tag{37}$$

which only proves the validity of the prescription given in (32).

6.4 Finding madhyama from manda-sphuta

The last section (Para (6)) of the manuscript deliniates the procedure for obtaining the *madhyama* from the *manda-sphuta*. In fact the author commences with the declaration:

एवं कर्णानयने ज्ञाते स्फुटतो मध्यमानयनं निरूप्यते –

Thus having known how to find the *karṇa*, the procedure for obtaining the mean from the true is now explained.

The karna that is being referred to above is the avises akarna given by (31). It is said that having known this karna K, it should be multiplied by sphutabhujāphala and divided by trijyā. The term sphutabhujāphala refers to $r_0 \sin(\theta_{ms} - \theta_m)$ and hence the prescrition given amounts to finding

$$\frac{K}{R}r_0\sin(\theta_{ms}-\theta_m).\tag{38}$$

From (37), this is the same as

$$\frac{R}{R_v} r_0 \sin(\theta_{ms} - \theta_m). \tag{39}$$

Now it can be easily seen that the above expression is the same as the RHS of (36c) but for the multiplying factor $trijy\bar{a} R$. Hence we have,

$$R\sin(\theta_0 - \theta_{ms}) = r_0 \sin(\theta_{ms} - \theta_m) \frac{R}{R_v}, \qquad (40)$$

Or, equivalently

$$\theta_0 - \theta_{ms} = (R\sin)^{-1} \left[r_0 \sin(\theta_{ms} - \theta_m) \frac{R}{R_v} \right], \tag{41}$$

from which the madhyama θ_0 can be obtained, since θ_{ms} and θ_m are already known. It may also be recalled that R_v can be obtained in terms of θ_{ms} and θ_m using (11). Thus an elegant procedure³⁴ for obtaining the madhyama from manda-sphuta has been described, which explains that title of the manuscript madhyamānayanaprakāra.

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 $^{^{34}}$ which does away with the arduous iterative procedure

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